

**WP4: Verifying the kinematic properties of STEREO/HI CMEs  
against in-situ CME observations and coronal sources**

## **In-situ comparison**

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The goal of Task 4.2 is to construct a comprehensive catalogue of ICMEs measured *in-situ* and to list their key parameters and the results of modeling based on *in-situ* data.

Relevant HELCTATS catalogs: **DATA CAT** and **ICMECAT**

All data to be made available until end of 2015

# Where we are now

## DATA CAT

- In-situ data of plasma and magnetic fields starting from January 2007. All in same format IDL sav.-files
- Venus Express, MESSENGER, Ulysses (UNIGRAZ), STEREO, WIND (UH)

## ICME CAT

- Clear ICMEs and magnetic flux ropes (MFR) (MESSENGER and VEX more strict as based only on magnetic field measurement)
- Around 1 AU (Wind, STEREO): 435 ICMEs
- Inner heliosphere (MESSENGER, VEX): 167 ICMEs

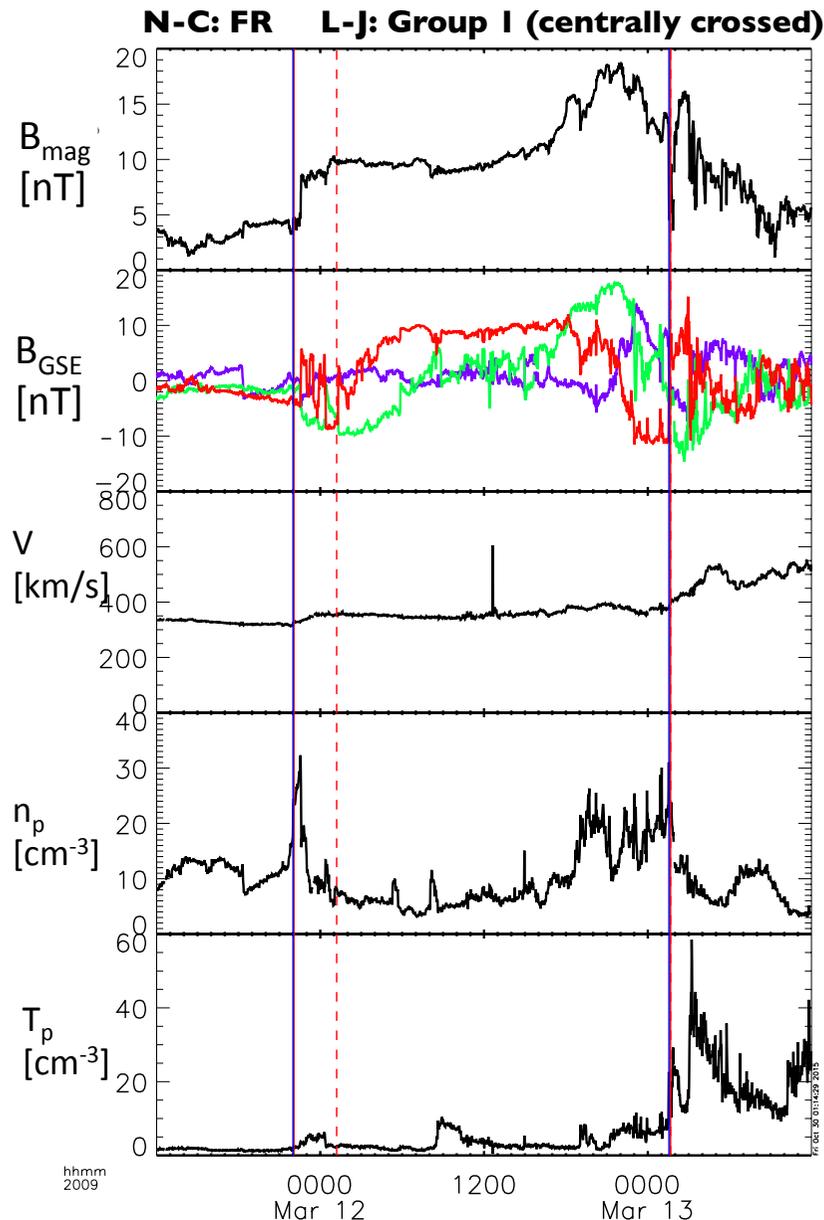
} 602 ICMEs

Spacecraft	Provider	Number of ICMEs	Start time	End time
Wind	T. Nieves-Chinchilla	135	Jan 2007	Dec 2013
STEREO-A	L. Jian	167	Jan 2007	Jun 2014
STEREO-B	L. Jian	133	Jan 2007	Dec 2013
VEX	S. Good	81	Jan 2007	Dec 2013
MESSENGER	S. Good	33	Jan 2007	May 2012
	R. Winslow	61	May 2011	Sep 2014

## Shocks

e.g., UH Heliospheric shock database

# Comparison of Nieves and Chinchilla (N-C) and Lan Jian (L-J) lists (WIND)



Line	Source	structure
blue solid	L-J	ICME
blue dashed	L-J	Obstacle (start)
red solid	N-C	ICME
red dashed	N-C	Obstacle

## N-C list

*ICME (solid)*

Start: shock or sheath signatures

End: Reverse shock or end of magnetic obstacle

*Magnetic Obstacle (dashed)*

Characterized as flux rope (FR, rotation), flux rope like (FRL, partial rotation), Ejecta (no rotation)

## L-J list

*ICME (solid)*

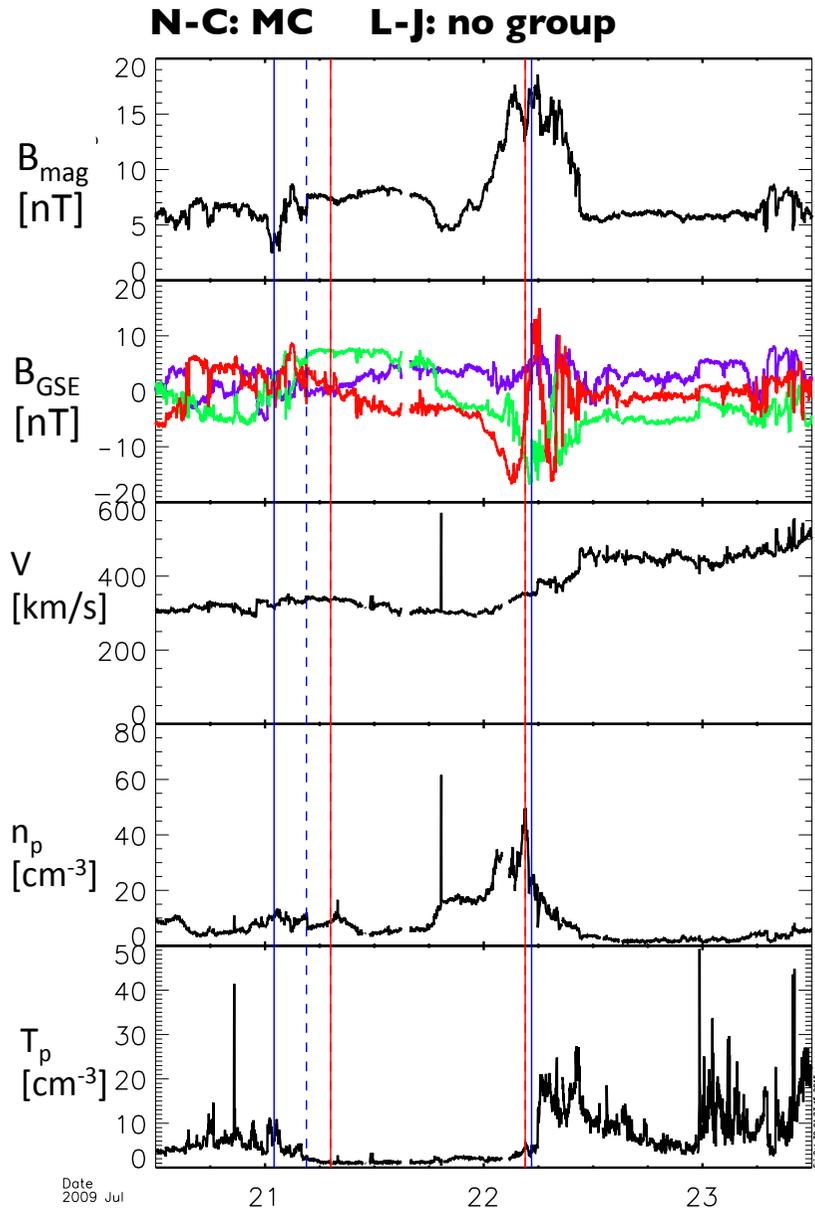
Start: Based on various signatures (Pt)

End: Based on various signatures (Pt)

*Magnetic Obstacle (dashed)*

End time same as for ICME. Start

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## L-J list

ICME (solid)

Start: Based on various signatures (Pt)

End: Based on various signatures (Pt)

Magnetic Obstacle (dashed)

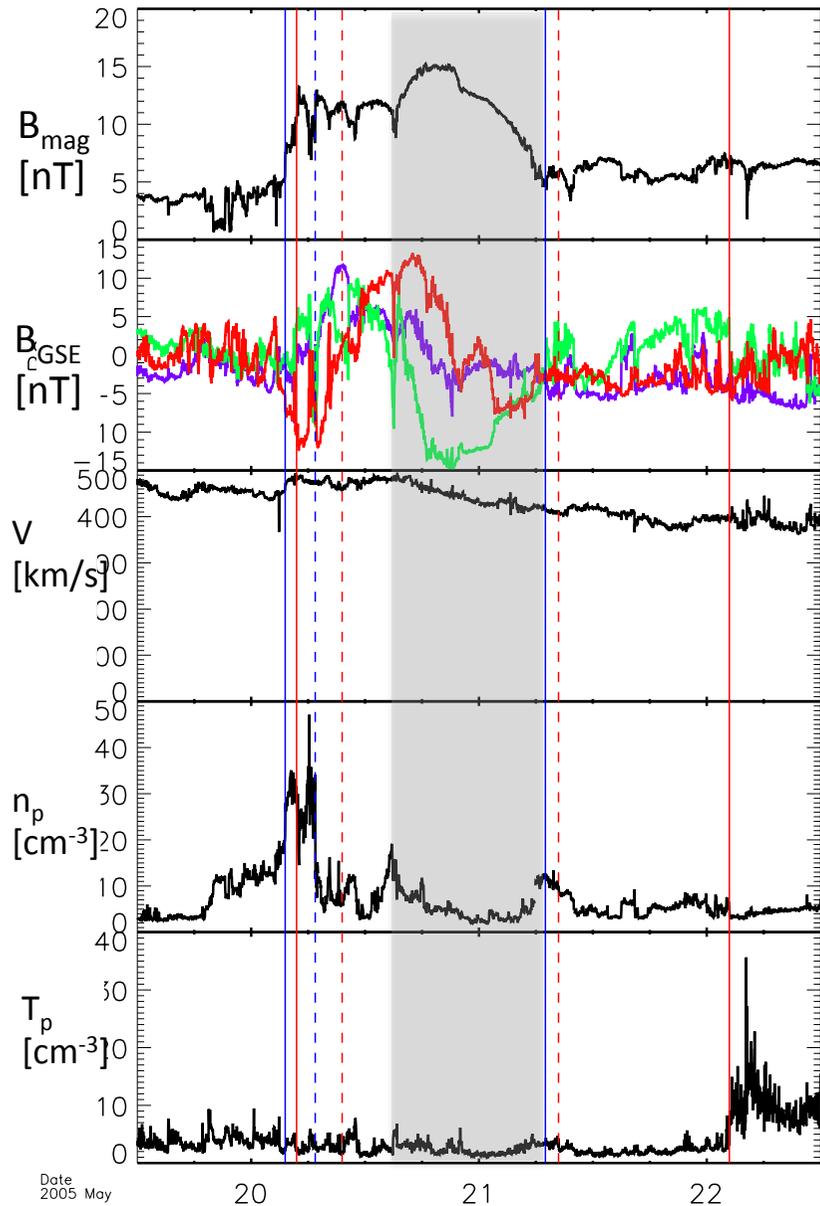
End time same as for ICME. Start

→ magnetic obstacle times for MVA and input to GS

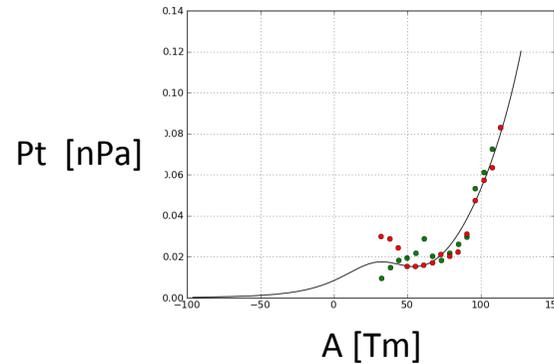
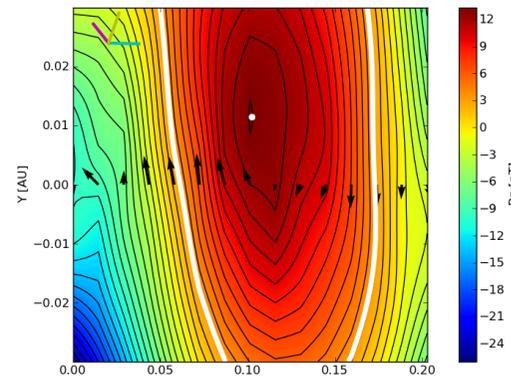
→ ICME – obstacle start times: Sheath values

# Comparison of Nieves and Chinchilla (N-C) and Lan Jian (L-J) lists (WIND)

N-C: FR    L-J: Group I (centrally crossed)



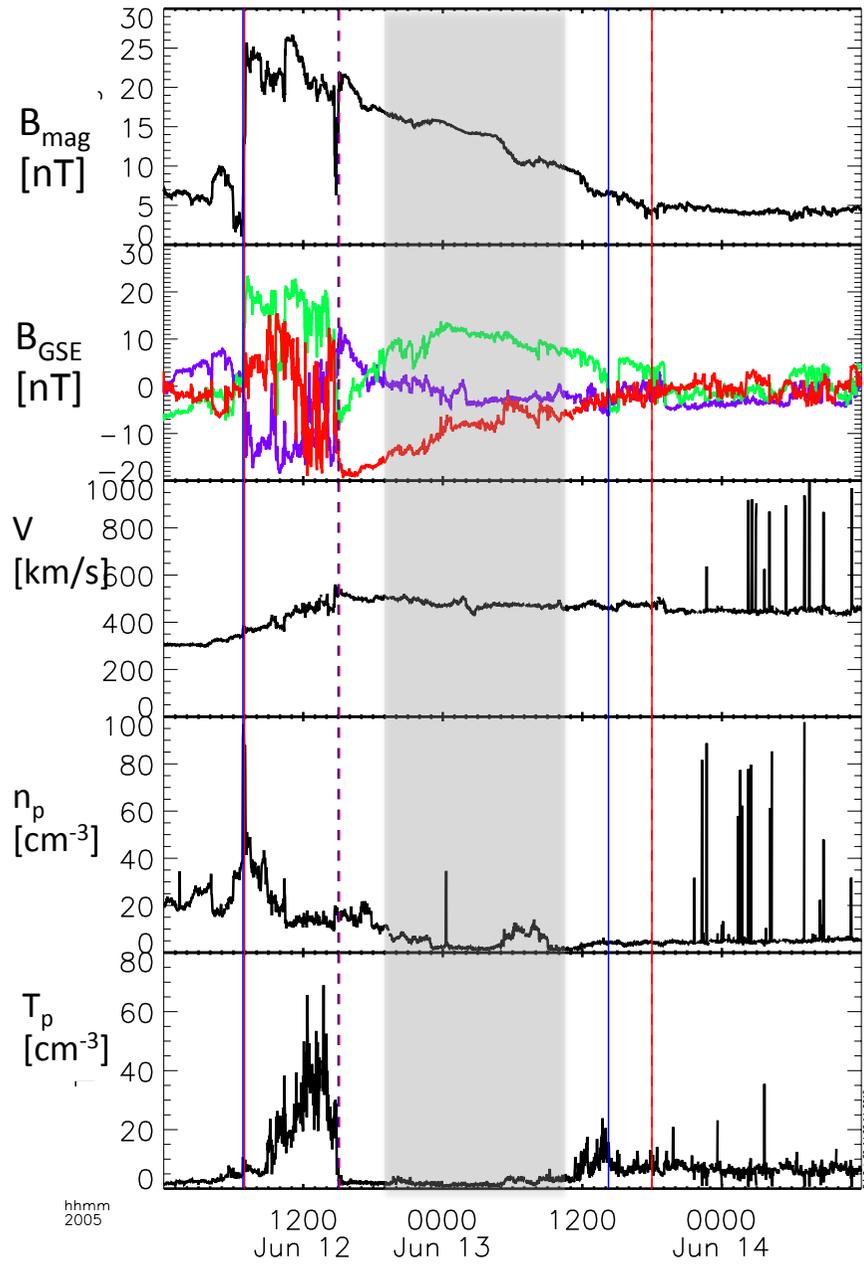
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red solid	N-C	ICME
red dashed	N-C	Obstacle



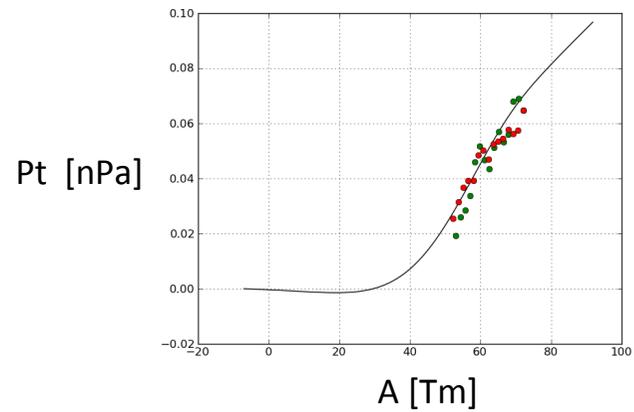
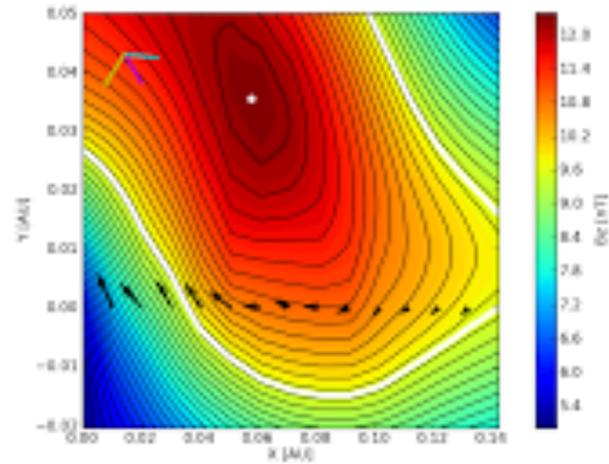
- Orientation
- Helicity
- Closes approach distance
- ( $V, B, n_p$  max etc.) automated

Details: Kilpua et al., 2013: <http://adsabs.harvard.edu/abs/2013AnGeo..31.1251K>

**N-C: FRL L-J: Group 3 (edge encounter)**



Line	Source	structure
blue solid	L-J	ICME
blue dashed	L-J	Obstacle (start)
red solid	N-C	ICME
red dashed	N-C	Obstacle



## **EGU 2016 Session:**

### **ST1.7**

Insights for Early Predictions of Magnetic and Dynamic Properties of Interplanetary Coronal Mass Ejections using Observations, Theory and Modeling

Convener: Emilia Kilpua

Co-Conveners: N. P. Savani, Spiros Patsourakos

