WP3: Deriving/cataloguing the kinetic properties of STEREO/HI CMEs based on geometrical and forward modelling

WP3.1: Geometrical modelling of STEREO/HI CMEs (STFC, UNIGRAZ, UGOE)









The objective of Task 3.1 is to perform geometrical modelling of the STEREO/HI CMEs, identified and catalogued in WP2, to derive their kinematic properties. This allows us to project back to the Sun and forward to specified solar system locations. From time-elongation maps (J-maps) generated from the HI data, the time-elongation profile of each CME will (where possible) be extracted and analysed using a range of single-spacecraft and stereoscopic geometric models (see Davies et al. 2012) to provide estimates of CME propagation speed, direction and potentially size; the STEREO/HI catalogue will be augmented with this information.









STEREO Heliospheric Imager HI-A: Ecliptic









Single-spacecraft timeelongation profile fitting

This requires a model of the transient's cross-sectional geometry:

• Fixed-Phi Fitting (FPF): point-like transient

 Harmonic Mean Fitting (HMF): circular transient fixed to Sun-centre.

 Self-Similar Expansion Fitting (SSE): generalisation of above into a single geometry



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HCME A 20081004 01	2008-10-04T20:49Z	А	50	95	NO	1	70	HCME A 20081004 01 PA070.dat	289	-25	205	19	2008-10-04T19:27Z
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For each tracked CME, we currently include in the catalogue:

- Filename of time-elongation profile;
- Radial speed;
- HEEQ longitude;
- Carrington longitude;
- HEEQ latitude.







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Geometrical modelling techniques

As well as kinematic properties, back-projected CME launch time/location, derived from the geometrical modelling, will be incorporated into the catalogue.

- This enables potential source signatures associated with CME onset to be identified (WP4.1).
- The in-situ observations of CMEs will be compared to their white-light counterparts (WP4.2).





